

Nuclear forces

Morten Hjorth-Jensen¹

National Superconducting Cyclotron Laboratory and Department of Physics and Astronomy, Michigan State University, East Lansing, MI 48824, USA¹

Spring 2016

© 2013-2016, Morten Hjorth-Jensen. Released under CC Attribution-NonCommercial 4.0 license

Components of the force and isospin

The nuclear forces are almost charge independent. If we assume they are, we can introduce a new quantum number which is conserved. For nucleons only, that is a proton and neutron, we can limit ourselves to two possible values which allow us to distinguish between the two particles. If we assign an isospin value of $\tau = 1/2$ for protons and neutrons (they belong to an isospin doublet, in the same way as we discussed the spin $1/2$ multiplet), we can define the neutron to have isospin projection $\tau_z = +1/2$ and a proton to have $\tau_z = -1/2$. These assignments are the standard choices in low-energy nuclear physics.

Phenomenology of nuclear forces

From Yukawa to Lattice QCD and Effective Field Theory

- Chadwick (1932) discovers the neutron and Heisenberg (1932) proposes the first Phenomenology (Isospin).
- Yukawa (1935) and his Meson Hypothesis
- Discovery of the pion in cosmic ray (1947) and in the Berkeley Cyclotron Lab (1948).
- Nobelprize awarded to Yukawa (1949). Rabi (1948) measures quadrupole moment of the deuteron.
- Taketani, Nakamura, Sasaki (1951): 3 ranges. One-Pion-Exchange (OPE): o.k.
- Multi-pion exchanges: Problems! Taketani, Machida, Onuma (1952);
- *Pion Theories* Brueckner, Watson (1953).

Phenomenology of nuclear forces

From Yukawa to Lattice QCD and Effective Field Theory

- Many pions = multi-pion resonances: $\sigma(600)$, $\rho(770)$, $\omega(782)$ etc. One-Boson-Exchange Model.
- Refined Meson Theories
- Sophisticated models for two-pion exchange:
 - Paris Potential (Lacombe et al., Phys. Rev. C **21**, 861 (1980))
 - Bonn potential (Machleidt et al., Phys. Rep. **149**, 1 (1987))

*Quark cluster models. Begin of effective field theory studies.

Phenomenology of nuclear forces

From Yukawa to Lattice QCD and Effective Field Theory

- 1990's
 - 1993-2001: High-precision NN potentials: Nijmegen I, II, '93, Reid93 (Stoks et al. 1994),
 - Argonne V18 (Wiringa et al, 1995), CD-Bonn (Machleidt et al. 1996 and 2001.
 - Advances in effective field theory: Weinberg (1990); Ordonez, Ray, van Kolck and many more.
- 3rd Millenium
 - Another "pion theory"; but now right: constrained by chiral symmetry. Three-body and higher-body forces appear naturally at a given order of the chiral expansion.

Nucleon-nucleon interaction from Lattice QCD, final confirmation of meson hypothesis of Yukawa? See for example Ishii *et al*, PRL 2007

Phenomenology of nuclear forces

Features of the Nucleon-Nucleon (NN) Force

The aim is to give you an overview over central features of the nucleon-nucleon interaction and how it is constructed, with both technical and theoretical approaches.

- The existence of the deuteron with $J^\pi = 1^+$ indicates that the force between protons and neutrons is attractive at least for the 3S_1 partial wave. Interference between Coulomb and nuclear scattering for the proton-proton partial wave 1S_0 shows that the NN force is attractive at least for the 1S_0 partial wave.
- It has a short range and strong intermediate attraction.
- Spin dependent, scattering lengths for triplet and singlet states are different,
- Spin-orbit force. Observation of large polarizations of scattered nucleons perpendicular to the plane of scattering.

Phenomenology of nuclear forces

- Strongly repulsive core. The s -wave phase shift becomes negative at ≈ 250 MeV implying that the singlet S has a hard core with range $0.4 - 0.5$ fm.
- Charge independence (almost). Two nucleons in a given two-body state always (almost) experience the same force. Modern interactions break charge and isospin symmetry lightly. That means that the pp , neutron-neutron and pn parts of the interaction will be different for the same quantum numbers.
- Non-central. There is a tensor force. First indications from the quadrupole moment of the deuteron pointing to an admixture in the ground state of both $l = 2$ (3D_1) and $l = 0$ (3S_1) orbital momenta.

Phenomenology of nuclear forces

Short Range Evidence

Comparison of the binding energies of ${}^2\text{H}$ (deuteron), ${}^3\text{H}$ (triton), ${}^4\text{He}$ (alpha - particle) show that the nuclear force is of finite range ($1 - 2$ fm) and very strong within that range. For nuclei with $A > 4$, the energy saturates: Volume and binding energies of nuclei are proportional to the mass number A (as we saw from exercise 1). Nuclei are also bound. The average distance between nucleons in nuclei is about 2 fm which must roughly correspond to the range of the attractive part.

Phenomenology of nuclear forces

Charge Dependence

- After correcting for the electromagnetic interaction, the forces between nucleons (pp , nn , or np) in the same state are almost the same.
- *Almost the same*: Charge-independence is slightly broken.
- Equality between the pp and nn forces: Charge symmetry.
- Equality between pp/nn force and np force: Charge independence.
- Better notation: Isospin symmetry, invariance under rotations in isospin

Phenomenology of nuclear forces

Charge Dependence, 1S_0 Scattering Lengths

Charge-symmetry breaking (CSB), after electromagnetic effects have been removed:

- $a_{pp} = -17.3 \pm 0.4$ fm
- $a_{nn} = -18.8 \pm 0.5$ fm. Note however discrepancy from nd breakup reactions resulting in $a_{nn} = -18.72 \pm 0.13 \pm 0.65$ fm and $\pi^- + d \rightarrow \gamma + 2n$ reactions giving $a_{nn} = -18.93 \pm 0.27 \pm 0.3$ fm.

Charge-independence breaking (CIB)

- $a_{pn} = -23.74 \pm 0.02$ fm

Symmetries of the Nucleon-Nucleon (NN) Force

- Translation invariance
- Galilean invariance
- Rotation invariance in space
- Space reflection invariance
- Time reversal invariance
- Invariance under the interchange of particle 1 and 2
- Almost isospin symmetry

A typical form of the nuclear force

Here we display a typical way to parametrize (non-relativistic expression) the nuclear two-body force in terms of some operators, the central part, the spin-spin part and the central force.

$$V(r) = \left\{ C_c + C_\sigma \sigma_1 \cdot \sigma_2 + C_T \left(1 + \frac{3}{m_\alpha r} + \frac{3}{(m_\alpha r)^2} \right) S_{12}(\hat{r}) + C_{SL} \left(\frac{1}{m_\alpha r} + \frac{1}{(m_\alpha r)^2} \right) \mathbf{L} \cdot \mathbf{S} \right\} \frac{e^{-m_\alpha r}}{m_\alpha r}$$

How do we derive such terms? (Note: no isospin dependence and that the above is an approximation)

Nuclear forces

To derive the above famous form of the nuclear force using field theoretical concepts, we will need some elements from relativistic quantum mechanics. These derivations will be given below. The material here gives some background to this. I know that many of you have not taken a course in quantum field theory, I hope however that you can see the basic ideas leading to the famous non-relativistic expressions for the nuclear force.

Furthermore, when we analyze nuclear data, we will actually try to explain properties like spectra, single-particle energies etc in terms of the various terms of the nuclear force. Moreover, many of you will hear about these terms at various talks, workshops, seminars etc. Then, it is good to have an idea of what people actually mean!!

Dramatis Personae

Baryons	Mass (MeV)	Mesons	Mass (MeV)
p, n	938.926	π	138.03
Λ	1116.0	η	548.8
Σ	1197.3	σ	≈ 550.0
Δ	1232.0	ρ	770
		ω	782.6
		δ	983.0
		K	495.8
		K^*	895.0

Components of the force and quantum numbers

But before we proceed, we will look into specific quantum numbers of the relative system and study expectation values of the various terms of

$$V(\mathbf{r}) = \left\{ C_c + C_\sigma \sigma_1 \cdot \sigma_2 + C_T \left(1 + \frac{3}{m_\alpha r} + \frac{3}{(m_\alpha r)^2} \right) S_{12}(\hat{r}) + C_{SL} \left(\frac{1}{m_\alpha r} + \frac{1}{(m_\alpha r)^2} \right) \mathbf{L} \cdot \mathbf{S} \right\} \frac{e^{-m_\alpha r}}{m_\alpha r}$$

Relative and CoM system, quantum numbers

When solving the scattering equation or solving the two-nucleon problem, it is convenient to rewrite the Schrodinger equation, due to the spherical symmetry of the Hamiltonian, in relative and center-of-mass coordinates. This will also define the quantum numbers of the relative and center-of-mass system and will aid us later in solving the so-called Lippman-Schwinger equation for the scattering problem.

We define the center-of-mass (CoM) momentum as

$$\mathbf{K} = \sum_{i=1}^A \mathbf{k}_i,$$

with $\hbar = c = 1$ the wave number $k_i = p_i$, with p_i the pertinent momentum of a single-particle state. We have also the relative momentum

$$\mathbf{k}_{ij} = \frac{1}{2}(\mathbf{k}_i - \mathbf{k}_j).$$

We will below skip the indices ij and simply write \mathbf{k}

Relative and CoM system, quantum numbers

In a similar fashion we can define the CoM coordinate

$$\mathbf{R} = \frac{1}{A} \sum_{i=1}^A \mathbf{r}_i,$$

and the relative distance

$$\mathbf{r}_{ij} = (\mathbf{r}_i - \mathbf{r}_j).$$

Relative and CoM system, quantum numbers

With the definitions

$$\mathbf{K} = \sum_{i=1}^A \mathbf{k}_i,$$

and

$$\mathbf{k}_{ij} = \frac{1}{2}(\mathbf{k}_i - \mathbf{k}_j).$$

we can rewrite the two-particle kinetic energy (note that we use $\hbar = c = 1$ as

$$\frac{\mathbf{k}_1^2}{2m_n} + \frac{\mathbf{k}_2^2}{2m_n} = \frac{\mathbf{k}^2}{m_n} + \frac{\mathbf{K}^2}{4m_n},$$

where m_n is the average of the proton and the neutron masses.

Relative and CoM system, quantum numbers

Since the two-nucleon interaction depends only on the relative distance, this means that we can separate Schroedinger's equation in an equation for the center-of-mass motion and one for the relative motion.

With an equation for the relative motion only and a separate one for the center-of-mass motion we need to redefine the two-body quantum numbers.

Previously we had a two-body state vector defined as $|(j_1 j_2) JM_J\rangle$ in a coupled basis. We will now define the quantum numbers for the relative motion. Here we need to define new orbital momenta (since these are the quantum numbers which change). We define

$$\hat{l}_1 + \hat{l}_2 = \hat{\lambda} = \hat{l} + \hat{L},$$

where \hat{l} is the orbital momentum associated with the relative motion and \hat{L} the corresponding one linked with the CoM. The total spin S is unchanged since it acts in a different space. We have thus that

$$\hat{I} = \hat{l} + \hat{L} + \hat{S}$$

Phenomenology of nuclear forces

The total two-nucleon state function has to be anti-symmetric. The total function contains a spatial part, a spin part and an isospin part. If isospin is conserved, this leads to in case we have an s -wave with spin $S = 0$ to an isospin two-body state with $T = 1$ since the spatial part is symmetric and the spin part is anti-symmetric.

Since the projections for T are $T_z = -1, 0, 1$, we can have a pp , an nn and a pn state.

For $l = 0$ and $S = 1$, a so-called triplet state, 3S_1 , we must have $T = 0$, meaning that we have only one state, a pn state. For other partial waves, the following table lists states up to f waves. We can systemize this in a table as follows, recalling that

$^{2S+1}l_J$	J	l	S	T	$ pp\rangle$	$ pn\rangle$	$ nn\rangle$
1S_0	0	0	0	1	yes	yes	yes
3S_1	1	0	1	0	no	yes	no
3P_0	0	1	1	1	yes	yes	yes
1P_1	1	1	0	0	no	yes	no
3P_1	1	1	1	1	yes	yes	yes
3P_2	2	1	1	1	yes	yes	yes
3D_1	1	2	1	0	no	yes	no

Components of the force and quantum numbers

The tensor force is given by

$$S_{12}(\hat{r}) = \frac{3}{r^2} (\sigma_1 \cdot \mathbf{r})(\sigma_2 \cdot \mathbf{r}) - \sigma_1 \cdot \sigma_2$$

where the Pauli matrices are defined as

$$\sigma_x = \begin{Bmatrix} 0 & 1 \\ 1 & 0 \end{Bmatrix},$$

$$\sigma_y = \begin{Bmatrix} 0 & -i \\ i & 0 \end{Bmatrix},$$

and

$$\sigma_z = \begin{Bmatrix} 1 & 0 \\ 0 & -1 \end{Bmatrix},$$

with the properties $\sigma = 2S$ (the spin of the system, being $1/2$ for nucleons), $\sigma_x^2 = \sigma_y^2 = \sigma_z^2 = \mathbf{1}$ and obeying the commutation and anti-commutation relations $\{\sigma_x, \sigma_y\} = 0$ $[\sigma_x, \sigma_y] = i\sigma_z$ etc.

Components of the force and quantum numbers

When we look at the expectation value of $\langle \sigma_1 \cdot \sigma_2 \rangle$, we can rewrite this expression in terms of the spin $S = s_1 + s_2$, resulting in

$$\langle \sigma_1 \cdot \sigma_2 \rangle = 2(S^2 - s_1^2 - s_2^2) = 2S(S+1) - 3,$$

where we $s_1 = s_2 = 1/2$ leading to

$$\begin{cases} \langle \sigma_1 \cdot \sigma_2 \rangle = 1 & \text{if } S = 1 \\ \langle \sigma_1 \cdot \sigma_2 \rangle = -3 & \text{if } S = 0 \end{cases}$$

Components of the force and quantum numbers

Similarly, the expectation value of the spin-orbit term is

$$\langle IS \rangle = \frac{1}{2} (J(J+1) - l(l+1) - S(S+1)),$$

which means that for s -waves with either $S = 0$ and thereby $J = 0$ or $S = 1$ and $J = 1$, the expectation value for the spin-orbit force is zero. With the above phenomenological model, the only contributions to the expectation value of the potential energy for s -waves stem from the central and the spin-spin components since the expectation value of the tensor force is also zero.

Components of the force and quantum numbers

For $s = 1/2$ spin values only for two nucleons, the expectation value of the tensor force operator is

l	$J+1$	J	$J-1$
$J+1$	$-\frac{2J(J+2)}{2J+1}$	0	$\frac{6\sqrt{J(J+1)}}{2J+1}$
J	0	2	0
$J-1$	$\frac{6\sqrt{J(J+1)}}{2J+1}$	0	$-\frac{2(2J+1)}{2J+1}$

We will derive these expressions after we have discussed the Wigner-Eckart theorem.

Components of the force and isospin

If we now add isospin to our simple V_4 interaction model, we end up with 8 operators, popularly dubbed V_8 interaction model. The explicit form reads

$$V(r) = \left\{ C_c + C_{\sigma} \sigma_1 \cdot \sigma_2 + C_T \left(1 + \frac{3}{m_{\alpha} r} + \frac{3}{(m_{\alpha} r)^2} \right) S_{12}(\hat{r}) \right. \\ \left. + C_{SL} \left(\frac{1}{m_{\alpha} r} + \frac{1}{(m_{\alpha} r)^2} \right) \mathbf{L} \cdot \mathbf{S} \right\} \frac{e^{-m_{\alpha} r}}{m_{\alpha} r} \\ + \left\{ C_{cr} + C_{\sigma r} \sigma_1 \cdot \sigma_2 + C_{Tr} \left(1 + \frac{3}{m_{\alpha} r} + \frac{3}{(m_{\alpha} r)^2} \right) S_{12}(\hat{r}) \right. \\ \left. + C_{SLr} \left(\frac{1}{m_{\alpha} r} + \frac{1}{(m_{\alpha} r)^2} \right) \mathbf{L} \cdot \mathbf{S} \right\} \tau_1 \cdot \tau_2 \frac{e^{-m_{\alpha} r}}{m_{\alpha} r}$$

Phenomenology of nuclear forces

References for Various Phenomenological Interactions

From 1950 till approximately 2000: One-Boson-Exchange (OBE) models dominate. These are models which typically include several low-mass mesons, that is with masses below 1 GeV. Potentials which are based upon the standard non-relativistic operator structure are called "Phenomenological Potentials" Some historically important examples are

- Gammel-Thaler potential (Phys. Rev. **107**, 291, 1339 (1957) and the
- Hamada-Johnston potential, Nucl. Phys. **34**, 382 (1962)), both with a hard core.
- Reid potential (Ann. Phys. (N.Y.) **50**, 411 (1968)), soft core.
- Argonne V_{14} potential (Wiringa et al., Phys. Rev. C **29**, 1207 (1984)) with 14 operators and the Argonne V_{18} potential (Wiringa et al., Phys. Rev. C **51**, 38 (1995)), uses 18 operators
- A good historical reference: R. Machleidt, Adv. Nucl. Phys.

Phenomenology of nuclear forces

The total two-nucleon state function has to be anti-symmetric. The total function contains a spatial part, a spin part and an isospin part. If isospin is conserved, this leads to in case we have an s -wave with spin $S = 0$ to an isospin two-body state with $T = 1$ since the spatial part is symmetric and the spin part is anti-symmetric. Since the projections for T are $T_z = -1, 0, 1$, we can have a pp , an nn and a pn state.

For $l = 0$ and $S = 1$, a so-called triplet state, 3S_1 , we must have $T = 0$, meaning that we have only one state, a pn state. For other partial waves, see exercises below.

Phenomenology of nuclear forces

Phenomenology of one-pion exchange

The one-pion exchange contribution (see derivation below), can be written as

$$V_{\pi}(r) = -\frac{f_{\pi}^2}{4\pi m_{\pi}^2} \tau_1 \cdot \tau_2 \frac{1}{3} \left\{ \sigma_1 \cdot \sigma_2 + \left(1 + \frac{3}{m_{\pi} r} + \frac{3}{(m_{\pi} r)^2} \right) S_{12}(\hat{r}) \right\} \frac{e^{-m_{\pi} r}}{m_{\pi} r}$$

Here the constant $f_{\pi}^2/4\pi \approx 0.08$ and the mass of the pion is $m_{\pi} \approx 140 \text{ MeV}/c^2$.

Phenomenology of nuclear forces

Let us look closer at specific partial waves for which one-pion exchange is applicable. If we have $S = 0$ and $T = 0$, the orbital momentum has to be an odd number in order for the total anti-symmetry to be obeyed. For $S = 0$, the tensor force component is zero, meaning that the only contribution is

$$V_{\pi}(r) = \frac{3f_{\pi}^2}{4\pi m_{\pi}^2} \frac{e^{-m_{\pi} r}}{m_{\pi} r},$$

since $\langle \sigma_1 \cdot \sigma_2 \rangle = -3$, that is we obtain a repulsive contribution to partial waves like 1P_0 .

Phenomenology of nuclear forces

Since $S = 0$ yields always a zero tensor force contribution, for the combination of $T = 1$ and then even l values, we get an attractive contribution

$$V_{\pi}(r) = -\frac{f_{\pi}^2}{4\pi m_{\pi}^2} \frac{e^{-m_{\pi} r}}{m_{\pi} r}.$$

With $S = 1$ and $T = 0$, l can only take even values in order to obey the anti-symmetry requirements and we get

$$V_{\pi}(r) = -\frac{f_{\pi}^2}{4\pi m_{\pi}^2} \left(1 + \left(1 + \frac{3}{m_{\pi} r} + \frac{3}{(m_{\pi} r)^2} \right) S_{12}(\hat{r}) \right) \frac{e^{-m_{\pi} r}}{m_{\pi} r},$$

while for $S = 1$ and $T = 1$, l can only take odd values, resulting in a repulsive contribution

$$V_{\pi}(r) = \frac{1}{3} \frac{f_{\pi}^2}{4\pi m_{\pi}^2} \left(1 + \left(1 + \frac{3}{m_{\pi} r} + \frac{3}{(m_{\pi} r)^2} \right) S_{12}(\hat{r}) \right) \frac{e^{-m_{\pi} r}}{m_{\pi} r}.$$

Phenomenology of nuclear forces

The central part of one-pion exchange interaction, arising from the spin-spin term, is thus attractive for s -waves and all even l values. For p -waves and all other odd values it is repulsive. However, its overall strength is weak. This is discussed further in one of exercises below.

Models for nuclear forces and derivation of non-relativistic expressions

To describe the interaction between the various baryons and mesons of the previous table we choose the following phenomenological lagrangians for spin 1/2 baryons

$$\mathcal{L}_{ps} = g^{ps} \bar{\Psi} \gamma^5 \Psi \phi^{(ps)},$$

$$\mathcal{L}_s = g^s \bar{\Psi} \Psi \phi^{(s)},$$

and

$$\mathcal{L}_v = g^v \bar{\Psi} \gamma_\mu \Psi \phi_\mu^{(v)} + g^t \bar{\Psi} \sigma^{\mu\nu} \Psi (\partial_\mu \phi_\nu^{(v)} - \partial_\nu \phi_\mu^{(v)}),$$

for pseudoscalar (ps), scalar (s) and vector (v) coupling, respectively. The factors g^v and g^t are the vector and tensor coupling constants, respectively.

Models for nuclear forces and derivation of non-relativistic expressions

For spin 1/2 baryons, the fields Ψ are expanded in terms of the Dirac spinors (positive energy solution shown here with $\bar{u}u = 1$)

$$u(k\sigma) = \sqrt{\frac{E(k) + m}{2m}} \begin{pmatrix} \chi \\ \frac{\sigma \mathbf{k}}{E(k) + m} \chi \end{pmatrix},$$

with χ the familiar Pauli spinor and $E(k) = \sqrt{m^2 + |\mathbf{k}|^2}$. The positive energy part of the field Ψ reads

$$\Psi(x) = \frac{1}{(2\pi)^{3/2}} \sum_{\mathbf{k}\sigma} u(k\sigma) \exp(-ikx) a_{\mathbf{k}\sigma},$$

with a being a fermion annihilation operator.

Models for nuclear forces and derivation of non-relativistic expressions

Expanding the free Dirac spinors in terms of $1/m$ (m is here the mass of the relevant baryon) results, to lowest order, in the familiar non-relativistic expressions for baryon-baryon potentials. The configuration space version of the interaction can be approximated as

$$V(\mathbf{r}) = \left\{ C_C^L + C_C^S + C_\sigma \sigma_1 \cdot \sigma_2 + C_T \left(1 + \frac{3}{m_\alpha r} + \frac{3}{(m_\alpha r)^2} \right) S_{12}(\hat{r}) \right. \\ \left. + C_{SL} \left(\frac{1}{m_\alpha r} + \frac{1}{(m_\alpha r)^2} \right) \mathbf{L} \cdot \mathbf{S} \right\} \frac{\exp(-m_\alpha r)}{m_\alpha r},$$

where m_α is the mass of the relevant meson and S_{12} is the familiar tensor term.

Models for nuclear forces and derivation of non-relativistic expressions

We derive now the non-relativistic one-pion exchange interaction. Here p_1, p'_1, p_2, p'_2 and $k = p_1 - p'_1$ denote four-momenta. The vertices are given by the pseudovector Lagrangian

$$\mathcal{L}_{p\nu} = \frac{f_\pi}{m_\pi} \bar{\psi} \gamma_5 \gamma_\mu \psi \partial^\mu \phi_\pi.$$

From the Feynman diagram rules we can write the two-body interaction as

$$V^{p\nu} = \frac{f_\pi^2}{m_\pi^2} \frac{\bar{u}(p'_1) \gamma_5 \gamma_\mu (p_1 - p'_1)^\mu u(p_1) \bar{u}(p'_2) \gamma_5 \gamma_\nu (p_2 - p'_2)^\nu u(p_2)}{(p_1 - p'_1)^2 - m_\pi^2}.$$

Models for nuclear forces and derivation of non-relativistic expressions

The factors $p_1 - p'_1 = p'_2 - p_2$ are both the four-momentum of the exchanged meson and come from the derivative of the meson field in the interaction Lagrangian. The Dirac spinors obey

$$\gamma_\mu p^\mu u(p) = m u(p) \\ \bar{u}(p) \gamma_\mu p^\mu = m \bar{u}(p).$$

Models for nuclear forces and derivation of non-relativistic expressions

Using these relations, together with $\{\gamma_5, \gamma_\mu\} = 0$, we find

$$\begin{aligned}\bar{u}(\mathbf{p}'_1)\gamma_5\gamma_\mu(\mathbf{p}_1 - \mathbf{p}'_1)^\mu u(\mathbf{p}_1) &= m\bar{u}(\mathbf{p}'_1)\gamma_5 u(\mathbf{p}_1) + \bar{u}(\mathbf{p}'_1)\gamma_\mu\mathbf{p}'_1{}^\mu\gamma_5 u(\mathbf{p}_1) \\ &= 2m\bar{u}(\mathbf{p}'_1)\gamma_5 u(\mathbf{p}_1)\end{aligned}$$

and

$$\bar{u}(\mathbf{p}'_2)\gamma_5\gamma_\mu(\mathbf{p}'_2 - \mathbf{p}_2)^\mu = -2m\bar{u}(\mathbf{p}'_2)\gamma_5 u(\mathbf{p}_2).$$

Models for nuclear forces and derivation of non-relativistic expressions

We get

$$V^{\rho\nu} = -\frac{f_\pi^2}{m_\pi^2} 4m^2 \frac{\bar{u}(\mathbf{p}'_1)\gamma_5 u(\mathbf{p}_1)\bar{u}(\mathbf{p}'_2)\gamma_5 u(\mathbf{p}_2)}{(\mathbf{p}_1 - \mathbf{p}'_1)^2 - m_\pi^2}.$$

By inserting expressions for the Dirac spinors, we find

$$\begin{aligned}\bar{u}(\mathbf{p}'_1)\gamma_5 u(\mathbf{p}_1) &= \sqrt{\frac{(E'_1 + m)(E_1 + m)}{4m^2}} \left(\chi^\dagger \quad -\frac{\boldsymbol{\sigma}_1 \cdot \mathbf{p}_1}{E'_1 + m} \chi^\dagger \right) \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \\ &\quad \times \begin{pmatrix} \chi \\ \frac{\boldsymbol{\sigma}_1 \cdot \mathbf{p}_1}{E_1 + m} \chi \end{pmatrix} \\ &= \sqrt{\frac{(E'_1 + m)(E_1 + m)}{4m^2}} \left(\frac{\boldsymbol{\sigma}_1 \cdot \mathbf{p}_1}{E_1 + m} - \frac{\boldsymbol{\sigma}_1 \cdot \mathbf{p}'_1}{E'_1 + m} \right)\end{aligned}$$

Models for nuclear forces and derivation of non-relativistic expressions

Similarly

$$\bar{u}(\mathbf{p}'_2)\gamma_5 u(\mathbf{p}_2) = \sqrt{\frac{(E'_2 + m)(E_2 + m)}{4m^2}} \left(\frac{\boldsymbol{\sigma}_2 \cdot \mathbf{p}_2}{E_2 + m} - \frac{\boldsymbol{\sigma}_2 \cdot \mathbf{p}'_2}{E'_2 + m} \right).$$

In the CM system we have $\mathbf{p}_2 = -\mathbf{p}_1$, $\mathbf{p}'_2 = -\mathbf{p}'_1$ and so $E_2 = E_1$, $E'_2 = E'_1$. We can then write down the relativistic contribution to the NN potential in the CM system:

$$\begin{aligned}V^{\rho\nu} &= -\frac{f_\pi^2}{m_\pi^2} 4m^2 \frac{1}{(\mathbf{p}_1 - \mathbf{p}'_1)^2 - m_\pi^2} \frac{(E_1 + m)(E'_1 + m)}{4m^2} \\ &\quad \times \left(\frac{\boldsymbol{\sigma}_1 \cdot \mathbf{p}_1}{E_1 + m} - \frac{\boldsymbol{\sigma}_1 \cdot \mathbf{p}'_1}{E'_1 + m} \right) \left(\frac{\boldsymbol{\sigma}_2 \cdot \mathbf{p}_1}{E_1 + m} - \frac{\boldsymbol{\sigma}_2 \cdot \mathbf{p}'_1}{E'_1 + m} \right).\end{aligned}$$

Models for nuclear forces and derivation of non-relativistic expressions

In the non-relativistic limit we have to lowest order

$$E_1 = \sqrt{\mathbf{p}_1^2 + m^2} \approx m \approx E'_1$$

and then $(\mathbf{p}_1 - \mathbf{p}'_1)^2 = -\mathbf{k}^2$, so we get for the contribution to the NN potential

$$\begin{aligned}V^{\rho\nu} &= -\frac{f_\pi^2}{m_\pi^2} 4m^2 \frac{1}{\mathbf{k}^2 + m_\pi^2} \frac{2m \cdot 2m}{4m^2} \frac{\boldsymbol{\sigma}_1 \cdot (\mathbf{p}_1 - \mathbf{p}'_1)}{2m} \cdot \frac{\boldsymbol{\sigma}_2 \cdot (\mathbf{p}_1 - \mathbf{p}'_1)}{2m} \\ &= -\frac{f_\pi^2}{m_\pi^2} \frac{(\boldsymbol{\sigma}_1 \cdot \mathbf{k})(\boldsymbol{\sigma}_2 \cdot \mathbf{k})}{\mathbf{k}^2 + m_\pi^2}.\end{aligned}$$

We have omitted exchange terms and the isospin term $\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2$.

Models for nuclear forces and derivation of non-relativistic expressions

We have

$$V^{\rho\nu}(\mathbf{k}) = -\frac{f_\pi^2}{m_\pi^2} \frac{(\boldsymbol{\sigma}_1 \cdot \mathbf{k})(\boldsymbol{\sigma}_2 \cdot \mathbf{k})}{\mathbf{k}^2 + m_\pi^2}.$$

In coordinate space we have

$$V^{\rho\nu}(\mathbf{r}) = \int \frac{d^3k}{(2\pi)^3} e^{i\mathbf{k}\cdot\mathbf{r}} V^{\rho\nu}(\mathbf{k})$$

resulting in

$$V^{\rho\nu}(\mathbf{r}) = -\frac{f_\pi^2}{m_\pi^2} \boldsymbol{\sigma}_1 \cdot \nabla \boldsymbol{\sigma}_2 \cdot \nabla \int \frac{d^3k}{(2\pi)^3} e^{i\mathbf{k}\cdot\mathbf{r}} \frac{1}{\mathbf{k}^2 + m_\pi^2}.$$

Models for nuclear forces and derivation of non-relativistic expressions

We obtain

$$V^{\rho\nu}(\mathbf{r}) = -\frac{f_\pi^2}{m_\pi^2} \boldsymbol{\sigma}_1 \cdot \nabla \boldsymbol{\sigma}_2 \cdot \nabla \frac{e^{-m_\pi r}}{r}.$$

Carrying out the differentiation of

$$V^{\rho\nu}(\mathbf{r}) = -\frac{f_\pi^2}{m_\pi^2} \boldsymbol{\sigma}_1 \cdot \nabla \boldsymbol{\sigma}_2 \cdot \nabla \frac{e^{-m_\pi r}}{r}.$$

we arrive at the famous one-pion exchange potential with central and tensor parts

$$V(\mathbf{r}) = -\frac{f_\pi^2}{m_\pi^2} \left\{ C_\sigma \boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2 + C_T \left(1 + \frac{3}{m_\alpha r} + \frac{3}{(m_\alpha r)^2} \right) S_{12}(r) \right\} \frac{\exp(-m_\pi r)}{m_\pi r}$$

For the full potential add the exchange part and the $\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2$ term as well. (Subtle point: there is a divergence which gets cancelled by using cutoffs) This leads to coefficients C_σ and C_T which are fitted to data.

Models for nuclear forces and derivation of non-relativistic expressions

When we perform similar non-relativistic expansions for scalar and vector mesons we obtain for the σ meson

$$V^\sigma = g_{\sigma NN}^2 \frac{1}{k^2 + m_\sigma^2} \left(-1 + \frac{q^2}{2M_N^2} - \frac{k^2}{8M_N^2} - \frac{\mathbf{LS}}{2M_N^2} \right).$$

We note an attractive central force and spin-orbit force. This term has an intermediate range. We have defined $1/2(p_1 + p_1') = q$. For the full potential add the exchange part and the isospin dependence as well.

Models for nuclear forces and derivation of non-relativistic expressions

We obtain for the ω meson

$$V^\omega = g_{\omega NN}^2 \frac{1}{k^2 + m_\omega^2} \left(1 - 3 \frac{\mathbf{LS}}{2M_N^2} \right).$$

We note a repulsive central force and an attractive spin-orbit force. This term has short range. For the full potential add the exchange part and the isospin dependence as well.

Models for nuclear forces and derivation of non-relativistic expressions

Finally for the ρ meson

$$V^\rho = g_{\rho NN}^2 \frac{k^2}{k^2 + m_\rho^2} \left(-2\sigma_1\sigma_2 + S_{12}(\hat{k}) \right) \tau_1\tau_2.$$

We note a tensor force with sign opposite to that of the pion. This term has short range. For the full potential add the exchange part and the isospin dependence as well.

Models for nuclear forces and derivation of non-relativistic expressions

- Can use a one-boson exchange picture to construct a nucleon-nucleon interaction a la QED
- Non-relativistic approximation yields amongst other things a spin-orbit force which is much stronger than in atoms.
- At large intermediate distances pion exchange dominates while pion resonances (other mesons) dominate at intermediate and short range
 - Potentials are parameterized to fit selected two-nucleon data, binding energies and scattering phase shifts.
- Nowadays, chiral perturbation theory gives an effective theory that allows a systematic expansion in terms of controllable parameters. Good basis for many-body physics

The Lippman-Schwinger equation for two-nucleon scattering

What follows now is a more technical discussion on how we can solve the two-nucleon problem. This will lead us to the so-called Lippman-Schwinger equation for the scattering problem and a rewrite of Schroedinger's equation in relative and center-of-mass coordinates.

Before we break down the Schroedinger equation into a partial wave decomposition, we derive now the so-called Lippman-Schwinger equation. We will do this in an operator form first. Thereafter, we rewrite it in terms of various quantum numbers such as relative momenta, orbital momenta etc. The Schroedinger equation in abstract vector representation is

$$(\hat{H}_0 + \hat{V})|\psi_n\rangle = E_n|\psi_n\rangle.$$

In our case for the two-body problem \hat{H}_0 is just the kinetic energy. We rewrite it as

$$(\hat{H}_0 - E_n)|\psi_n\rangle = -\hat{V}|\psi_n\rangle.$$

The Lippman-Schwinger equation for two-nucleon scattering

The equation

$$|\psi_n\rangle = \frac{1}{(E_n - \hat{H}_0)} \hat{V}|\psi_n\rangle,$$

is normally solved in an iterative fashion. We assume first that

$$|\psi_n\rangle = |\phi_n\rangle,$$

where $|\phi_n\rangle$ are the eigenfunctions of

$$\hat{H}_0|\phi_n\rangle = \omega_n|\phi_n\rangle$$

the so-called unperturbed problem. In our case, these will simply be the kinetic energies of the relative motion.

The Lippman-Schwinger equation for two-nucleon scattering

Inserting $|\phi_n\rangle$ on the right-hand side of

$$|\psi_n\rangle = \frac{1}{(E_n - \hat{H}_0)} \hat{V} |\psi_n\rangle,$$

yields

$$|\psi_n\rangle = |\phi_n\rangle + \frac{1}{(E_n - \hat{H}_0)} \hat{V} |\phi_n\rangle,$$

as our first iteration. Reinserting again gives

$$|\psi_n\rangle = |\phi_n\rangle + \frac{1}{(E_n - \hat{H}_0)} \hat{V} |\phi_n\rangle + \frac{1}{(E_n - \hat{H}_0)} \hat{V} \frac{1}{(E_n - \hat{H}_0)} \hat{V} |\phi_n\rangle,$$

and continuing we obtain

$$|\psi_n\rangle = \sum_{i=0}^{\infty} \left[\frac{1}{(E_n - \hat{H}_0)} \hat{V} \right]^i |\phi_n\rangle.$$

The Lippman-Schwinger equation for two-nucleon scattering

It is easy to see that

$$|\psi_n\rangle = \sum_{i=0}^{\infty} \left[\frac{1}{(E_n - \hat{H}_0)} \hat{V} \right]^i |\phi_n\rangle,$$

can be rewritten as

$$|\psi_n\rangle = |\phi_n\rangle + \frac{1}{(E_n - \hat{H}_0)} \hat{V} \left(1 + \frac{1}{(E_n - \hat{H}_0)} \hat{V} + \frac{1}{(E_n - \hat{H}_0)} \hat{V} \frac{1}{(E_n - \hat{H}_0)} \hat{V} \right) |\phi_n\rangle,$$

which we rewrite as

$$|\psi_n\rangle = |\phi_n\rangle + \frac{1}{(E_n - \hat{H}_0)} \hat{V} |\psi_n\rangle.$$

The Lippman-Schwinger equation for two-nucleon scattering

In operator form we have thus

$$|\psi_n\rangle = |\phi_n\rangle + \frac{1}{(E_n - \hat{H}_0)} \hat{V} |\psi_n\rangle.$$

We multiply from the left with $\langle \phi_m |$ and obtain

$$\langle \phi_m | \hat{V} |\psi_n\rangle = \langle \phi_m | \hat{V} |\phi_n\rangle + \langle \phi_m | \hat{V} \frac{1}{(E_n - \hat{H}_0)} \hat{V} |\psi_n\rangle.$$

We define thereafter the so-called T -matrix as

$$\langle \phi_m | \hat{T} |\phi_n\rangle = \langle \phi_m | \hat{V} |\psi_n\rangle.$$

We can rewrite our equation as

$$\langle \phi_m | \hat{T} |\phi_n\rangle = \langle \phi_m | \hat{V} |\phi_n\rangle + \langle \phi_m | \hat{V} \frac{1}{(E_n - \hat{H}_0)} \hat{T} |\phi_n\rangle.$$

The Lippman-Schwinger equation for two-nucleon scattering

The equation

$$\langle \phi_m | \hat{T} |\phi_n\rangle = \langle \phi_m | \hat{V} |\phi_n\rangle + \langle \phi_m | \hat{V} \frac{1}{(E_n - \hat{H}_0)} \hat{T} |\phi_n\rangle,$$

is called the Lippman-Schwinger equation. Inserting the completeness relation

$$\mathbf{1} = \sum_n |\phi_n\rangle \langle \phi_n|, \quad \langle \phi_n | \phi_{n'}\rangle = \delta_{n,n'}$$

we have

$$\langle \phi_m | \hat{T} |\phi_n\rangle = \langle \phi_m | \hat{V} |\phi_n\rangle + \sum_k \langle \phi_m | \hat{V} |\phi_k\rangle \frac{1}{(E_n - \omega_k)} \langle \phi_k | \hat{T} |\phi_n\rangle,$$

which is (when we specify the state $|\phi_n\rangle$) an integral equation that can actually be solved by matrix inversion easily! The unknown quantity is the T -matrix.

The Lippman-Schwinger equation for two-nucleon scattering

Now we wish to introduce a partial wave decomposition in order to solve the Lippman-Schwinger equation. With a partial wave decomposition we can reduce a three-dimensional integral equation to a one-dimensional one.

Let us continue with our Schrodinger equation in the abstract vector representation

$$(T + V) |\psi_n\rangle = E_n |\psi_n\rangle$$

Here T is the kinetic energy operator and V is the potential operator. The eigenstates form a complete orthonormal set according to

$$\mathbf{1} = \sum_n |\psi_n\rangle \langle \psi_n|, \quad \langle \psi_n | \psi_{n'}\rangle = \delta_{n,n'}$$

The Lippman-Schwinger equation for two-nucleon scattering

The most commonly used representations are the coordinate and the momentum space representations. They define the completeness relations

$$\mathbf{1} = \int d\mathbf{r} |\mathbf{r}\rangle \langle \mathbf{r}|, \quad \langle \mathbf{r} | \mathbf{r}'\rangle = \delta(\mathbf{r} - \mathbf{r}')$$

$$\mathbf{1} = \int d\mathbf{k} |\mathbf{k}\rangle \langle \mathbf{k}|, \quad \langle \mathbf{k} | \mathbf{k}'\rangle = \delta(\mathbf{k} - \mathbf{k}')$$

Here the basis states in both \mathbf{r} - and \mathbf{k} -space are dirac-delta function normalized. From this it follows that the plane-wave states are given by,

$$\langle \mathbf{r} | \mathbf{k}\rangle = \left(\frac{1}{2\pi} \right)^{3/2} \exp(i\mathbf{k} \cdot \mathbf{r})$$

which is a transformation function defining the mapping from the abstract $|\mathbf{k}\rangle$ to the abstract $|\mathbf{r}\rangle$ space.

The Lippman-Schwinger equation for two-nucleon scattering

That the r-space basis states are delta-function normalized follows from

$$\delta(\mathbf{r}-\mathbf{r}') = \langle \mathbf{r} | \mathbf{r}' \rangle = \langle \mathbf{r} | \mathbf{1} | \mathbf{r}' \rangle = \int d\mathbf{k} \langle \mathbf{r} | \mathbf{k} \rangle \langle \mathbf{k} | \mathbf{r}' \rangle = \left(\frac{1}{2\pi} \right)^3 \int d\mathbf{k} e^{i\mathbf{k}(\mathbf{r}-\mathbf{r}')}$$

and the same for the momentum space basis states,

$$\delta(\mathbf{k}-\mathbf{k}') = \langle \mathbf{k} | \mathbf{k}' \rangle = \langle \mathbf{k} | \mathbf{1} | \mathbf{k}' \rangle = \int d\mathbf{r} \langle \mathbf{k} | \mathbf{r} \rangle \langle \mathbf{r} | \mathbf{k}' \rangle = \left(\frac{1}{2\pi} \right)^3 \int d\mathbf{r} e^{i\mathbf{r}(\mathbf{k}-\mathbf{k}')}$$

The Lippman-Schwinger equation for two-nucleon scattering

Projecting on momentum states, we obtain the momentum space Schroedinger equation as

$$\frac{\hbar^2}{2\mu} k^2 \psi_n(\mathbf{k}) + \int d\mathbf{k}' V(\mathbf{k}, \mathbf{k}') \psi_n(\mathbf{k}') = E_n \psi_n(\mathbf{k}) \quad (1)$$

Here the notation $\psi_n(\mathbf{k}) = \langle \mathbf{k} | \psi_n \rangle$ and $\langle \mathbf{k} | V | \mathbf{k}' \rangle = V(\mathbf{k}, \mathbf{k}')$ has been introduced. The potential in momentum space is given by a double Fourier-transform of the potential in coordinate space, i.e.

$$V(\mathbf{k}, \mathbf{k}') = \left(\frac{1}{2\pi} \right)^3 \int d\mathbf{r} \int d\mathbf{r}' \exp -i\mathbf{k}\mathbf{r} V(\mathbf{r}, \mathbf{r}') \exp i\mathbf{k}'\mathbf{r}'$$

The Lippman-Schwinger equation for two-nucleon scattering

Here it is assumed that the potential interaction does not contain any spin dependence. Instead of a differential equation in coordinate space, the Schroedinger equation becomes an integral equation in momentum space. This has many tractable features. Firstly, most realistic nucleon-nucleon interactions derived from field-theory are given explicitly in momentum space. Secondly, the boundary conditions imposed on the differential equation in coordinate space are automatically built into the integral equation. And last, but not least, integral equations are easy to numerically implement, and convergence is obtained by just increasing the number of integration points. Instead of solving the three-dimensional integral equation, an infinite set of 1-dimensional equations can be obtained via a partial wave expansion.

The Lippman-Schwinger equation for two-nucleon scattering

The wave function $\psi_n(\mathbf{k})$ can be expanded in a complete set of spherical harmonics, that is

$$\psi_n(\mathbf{k}) = \sum_{lm} \psi_{nlm}(k) Y_{lm}(\hat{\mathbf{k}}) \quad \psi_{nlm}(k) = \int d\hat{\mathbf{k}} Y_{lm}^*(\hat{\mathbf{k}}) \psi_n(\mathbf{k}), \quad (2)$$

By inserting equation (2) in equation (1), and projecting from the left $Y_{lm}(\hat{\mathbf{k}})$, the three-dimensional Schroedinger equation (1) is reduced to an infinite set of 1-dimensional angular momentum coupled integral equations,

$$\left(\frac{\hbar^2}{2\mu} k^2 - E_{nlm} \right) \psi_{nlm}(k) = - \sum_{l'm'} \int_0^\infty dk' k'^2 V_{lm,l'm'}(k, k') \psi_{nl'm'}(k') \quad (3)$$

where the angular momentum projected potential takes the form,

$$V_{lm,l'm'}(k, k') = \int d\hat{\mathbf{k}} \int d\hat{\mathbf{k}}' Y_{lm}^*(\hat{\mathbf{k}}) V(\mathbf{k}\hat{\mathbf{k}}) Y_{l'm'}(\hat{\mathbf{k}}') \quad (4)$$

The Lippman-Schwinger equation for two-nucleon scattering

The potential is often given in position space. It is then convenient to establish the connection between $V_{lm,l'm'}(k, k')$ and $V_{lm,l'm'}(r, r')$. Inserting the completeness relation for the position quantum numbers in equation (4) results in

$$V = \int d\mathbf{r} \int d\mathbf{r}' \left\{ \int d\hat{\mathbf{k}} Y_{lm}^*(\hat{\mathbf{k}}) \langle \mathbf{k} | \mathbf{r} \rangle \right\} \langle \mathbf{r} | V | \mathbf{r}' \rangle \left\{ \int d\hat{\mathbf{k}}' Y_{l'm'}(\hat{\mathbf{k}}') \langle \mathbf{r}' | \mathbf{k}' \rangle \right\} \quad (5)$$

The Lippman-Schwinger equation for two-nucleon scattering

Since the plane waves depend only on the absolute values of position and momentum, $|\mathbf{k}|$ and $|\mathbf{r}|$, and the angle between them, θ_{kr} , they may be expanded in terms of bipolar harmonics of zero rank, i.e.

$$\exp(i\mathbf{k} \cdot \mathbf{r}) = 4\pi \sum_{l=0}^{\infty} i^l j_l(kr) \left(Y_l(\hat{\mathbf{k}}) \cdot Y_l(\hat{\mathbf{r}}) \right) = \sum_{l=0}^{\infty} (2l+1) i^l j_l(kr) P_l(\cos \theta_{kr})$$

where the addition theorem for spherical harmonics has been used in order to write the expansion in terms of Legendre polynomials. The spherical Bessel functions, $j_l(z)$, are given in terms of Bessel functions of the first kind with half integer orders,

$$j_l(z) = \sqrt{\frac{\pi}{2z}} J_{l+1/2}(z).$$

The Lippman-Schwinger equation for two-nucleon scattering

Inserting the plane-wave expansion into the brackets of equation (5) yields,

$$\int d\hat{k} Y_{lm}^*(\hat{k}) \langle \mathbf{k} | \mathbf{r} \rangle = \left(\frac{1}{2\pi} \right)^{3/2} 4\pi i^{-l} j_l(kr) Y_{lm}^*(\hat{r}),$$

$$\int d\hat{k}' Y_{lm}(\hat{k}') \langle \mathbf{r}' | \mathbf{k}' \rangle = \left(\frac{1}{2\pi} \right)^{3/2} 4\pi i^{l'} j_{l'}(k'r') Y_{l'm'}(\hat{r}').$$

The Lippman-Schwinger equation for two-nucleon scattering

The connection between the momentum- and position space angular momentum projected potentials are then given,

$$V_{lm,l'm'}(k, k') = \frac{2}{\pi} i^{l'-l} \int_0^\infty dr r^2 \int_0^\infty dr' r'^2 j_l(kr) V_{lm,l'm'}(r, r') j_{l'}(k'r')$$

which is known as a double Fourier-Bessel transform. The position space angular momentum projected potential is given by

$$V_{lm,l'm'}(r, r') = \int d\hat{r} \int d\hat{r}' Y_{lm}^*(\hat{r}) V(\mathbf{r}, \mathbf{r}') Y_{l'm'}(\hat{r}').$$

The Lippman-Schwinger equation for two-nucleon scattering

No assumptions of locality/non-locality and deformation of the interaction has so far been made, and the result in equation (62) is general. In position space the Schroedinger equation takes form of an integro-differential equation in case of a non-local interaction, in momentum space the Schroedinger equation is an ordinary integral equation of the Fredholm type, see equation (3). This is a further advantage of the momentum space approach as compared to the standard position space approach. If we assume that the interaction is of local character, i.e.

$$\langle \mathbf{r} | V | \mathbf{r}' \rangle = V(\mathbf{r}) \delta(\mathbf{r} - \mathbf{r}') = V(\mathbf{r}) \frac{\delta(r - r')}{r^2} \delta(\cos \theta - \cos \theta') \delta(\varphi - \varphi'),$$

then equation (62) reduces to

$$V_{lm,l'm'}(r, r') = \frac{\delta(r - r')}{r^2} \int d\hat{r} Y_{lm}^*(\hat{r}) V(\mathbf{r}) Y_{l'm'}(\hat{r}), \quad (6)$$

The Lippman-Schwinger equation for two-nucleon scattering

and equation (62) reduces to

$$V_{lm,l'm'}(k, k') = \frac{2}{\pi} i^{l'-l} \int_0^\infty dr r^2 j_l(kr) V_{lm,l'm'}(r) j_{l'}(k'r) \quad (7)$$

where

$$V_{lm,l'm'}(r) = \int d\hat{r} Y_{lm}^*(\hat{r}) V(\mathbf{r}) Y_{l'm'}(\hat{r}), \quad (8)$$

The Lippman-Schwinger equation for two-nucleon scattering

In the case that the interaction is central, $V(\mathbf{r}) = V(r)$, then

$$V_{lm,l'm'}(r) = V(r) \int d\hat{r} Y_{lm}^*(\hat{r}) Y_{l'm'}(\hat{r}) = V(r) \delta_{l,l'} \delta_{m,m'}, \quad (9)$$

and

$$V_{lm,l'm'}(k, k') = \frac{2}{\pi} \int_0^\infty dr r^2 j_l(kr) V(r) j_{l'}(k'r) \delta_{l,l'} \delta_{m,m'} = V_l(k, k') \delta_{l,l'} \delta_{m,m'}, \quad (10)$$

where the momentum space representation of the interaction finally reads,

$$V_l(k, k') = \frac{2}{\pi} \int_0^\infty dr r^2 j_l(kr) V(r) j_l(k'r). \quad (11)$$

The Lippman-Schwinger equation for two-nucleon scattering

For a local and spherical symmetric potential, the coupled momentum space Schroedinger equations given in equation (3) decouples in angular momentum, giving

$$\frac{\hbar^2}{2\mu} k^2 \psi_{nl}(k) + \int_0^\infty dk' k'^2 V_l(k, k') \psi_{nl}(k') = E_n \psi_{nl}(k) \quad (12)$$

Where we have written $\psi_{nl}(k) = \psi_{nlm}(k)$, since the equation becomes independent of the projection m for spherical symmetric interactions. The momentum space wave functions $\psi_{nl}(k)$ defines a complete orthogonal set of functions, which spans the space of functions with a positive finite Euclidean norm (also called l^2 -norm), $\sqrt{\langle \psi_n | \psi_n \rangle}$, which is a Hilbert space. The corresponding normalized wave function in coordinate space is given by the Fourier-Bessel transform

$$\phi_{nl}(r) = \sqrt{\frac{2}{\pi}} \int dk k^2 j_l(kr) \psi_{nl}(k)$$

The Lippman-Schwinger equation for two-nucleon scattering

We will thus assume that the interaction is spherically symmetric and use the partial wave expansion of the plane waves in terms of spherical harmonics. This means that we can separate the radial part of the wave function from its angular dependence. The wave function of the relative motion is described in terms of plane waves as

$$\exp(i\mathbf{k}\mathbf{r}) = \langle \mathbf{r} | \mathbf{k} \rangle = 4\pi \sum_{lm} i^l j_l(kr) Y_{lm}^*(\hat{\mathbf{k}}) Y_{lm}(\hat{\mathbf{r}}),$$

where j_l is a spherical Bessel function and Y_{lm} the spherical harmonics.

The Lippman-Schwinger equation for two-nucleon scattering

In terms of the relative and center-of-mass momenta \mathbf{k} and \mathbf{K} , the potential in momentum space is related to the nonlocal operator $V(\mathbf{r}, \mathbf{r}')$ by

$$\langle \mathbf{k}' \mathbf{K}' | V | \mathbf{k} \mathbf{K} \rangle = \int d\mathbf{r} d\mathbf{r}' \exp(-i\mathbf{k}'\mathbf{r}') V(\mathbf{r}', \mathbf{r}) \exp i\mathbf{k}\mathbf{r} \delta(\mathbf{K}, \mathbf{K}').$$

We will assume that the interaction is spherically symmetric. Can separate the radial part of the wave function from its angular dependence. The wave function of the relative motion is described in terms of plane waves as

$$\exp(i\mathbf{k}\mathbf{r}) = \langle \mathbf{r} | \mathbf{k} \rangle = 4\pi \sum_{lm} i^l j_l(kr) Y_{lm}^*(\hat{\mathbf{k}}) Y_{lm}(\hat{\mathbf{r}}),$$

where j_l is a spherical Bessel function and Y_{lm} the spherical harmonic.

The Lippman-Schwinger equation for two-nucleon scattering

This partial wave basis is useful for defining the operator for the nucleon-nucleon interaction, which is symmetric with respect to rotations, parity and isospin transformations. These symmetries imply that the interaction is diagonal with respect to the quantum numbers of total relative angular momentum \mathcal{J} , spin S and isospin T (we skip isospin for the moment). Using the above plane wave expansion, and coupling to final \mathcal{J} and S and T we get

$$\langle \mathbf{k}' | V | \mathbf{k} \rangle = (4\pi)^2 \sum_{STI'l'm_l m_l' \mathcal{J}} i^{l+l'} Y_{lm}^*(\hat{\mathbf{k}}) Y_{l'm'}(\hat{\mathbf{k}}')$$

$$\langle l m_l S m_S | \mathcal{J} M \rangle \langle l' m_l' S m_S | \mathcal{J} M \rangle \langle k' l' S \mathcal{J} M | V | k l S \mathcal{J} M \rangle,$$

where we have defined

$$\langle k' l' S \mathcal{J} M | V | k l S \mathcal{J} M \rangle = \int j_{l'}(k' r') \langle l' S \mathcal{J} M | V(r', r) | l S \mathcal{J} M \rangle j_l(kr) r^2 dr r'^2$$

We have omitted the momentum of the center-of-mass motion \mathbf{K} and the corresponding orbital momentum L , since the interaction is

The Lippman-Schwinger equation for two-nucleon scattering

We wrote the Lippman-Schwinger equation as

$$\langle \phi_m | \hat{T} | \phi_n \rangle = \langle \phi_m | \hat{V} | \phi_n \rangle + \sum_k \langle \phi_m | \hat{V} | \phi_k \rangle \frac{1}{(E_n - \omega_k)} \langle \phi_k | \hat{T} | \phi_n \rangle.$$

How do we rewrite it in a partial wave expansion with momenta k ?

The Lippman-Schwinger equation for two-nucleon scattering

The general structure of the T -matrix in partial waves is

$$T_{ll'}^\alpha(kk'K\omega) = V_{ll'}^\alpha(kk') + \frac{2}{\pi} \sum_{l'' m_l'' M_S} \int_0^\infty dq (\langle l'' m_l'' S m_S | \mathcal{J} M \rangle)^2 \frac{Y_{l'' m_l''}^*(\hat{\mathbf{q}}) Y_{l'' m_l''}(\hat{\mathbf{q}}) V_{ll''}^\alpha(kq) T_{l'' l'}^\alpha(kk')}{\omega - H_0} \quad (13)$$

The Lippman-Schwinger equation for two-nucleon scattering

The shorthand notation

$$T_{ll'}^\alpha(kk'K\omega) = \langle k K l L \mathcal{J} S | T(\omega) | k' K' l' L \mathcal{J} S \rangle,$$

denotes the T -matrix with momenta k and k' and orbital momenta l and l' of the relative motion, and K is the corresponding momentum of the center-of-mass motion. Further, L , \mathcal{J} , S and T are the orbital momentum of the center-of-mass motion, the total angular momentum, spin and isospin, respectively. Due to the nuclear tensor force, the interaction is not diagonal in $l'l'$.

The Lippman-Schwinger equation for two-nucleon scattering

Using the orthogonality properties of the Clebsch-Gordan coefficients and the spherical harmonics, we obtain the well-known one-dimensional angle independent integral equation

$$T_{l'l'}^\alpha(kk'K\omega) = V_{l'l'}^\alpha(kk') + \frac{2}{\pi} \sum_{l''} \int_0^\infty dq q^2 \frac{V_{l''}^\alpha(kq) T_{l''l'}^\alpha(qk'K\omega)}{\omega - H_0}$$

Inserting the denominator we arrive at

$$\hat{T}_{l'l'}^\alpha(kk'K) = \hat{V}_{l'l'}^\alpha(kk') + \frac{2}{\pi} \sum_{l''} \int_0^\infty dq q^2 \hat{V}_{l''}^\alpha(kq) \frac{1}{k^2 - q^2 + i\epsilon} \hat{T}_{l''l'}^\alpha(qk'K)$$

The Lippman-Schwinger equation for two-nucleon scattering

To parameterize the nucleon-nucleon interaction we solve the Lippman-Schwinger equation

$$T_{l'l'}^\alpha(kk'K) = V_{l'l'}^\alpha(kk') + \frac{2}{\pi} \sum_{l''} \int_0^\infty dq q^2 V_{l''}^\alpha(kq) \frac{1}{k^2 - q^2 + i\epsilon} T_{l''l'}^\alpha(qk'K)$$

The shorthand notation

$$T(\hat{V})_{l'l'}^\alpha(kk'K\omega) = \langle kKlL\mathcal{J}S | T(\omega) | k'K'l'\mathcal{J}S \rangle,$$

denotes the $T(V)$ -matrix with momenta k and k' and orbital momenta l and l' of the relative motion, and K is the corresponding momentum of the center-of-mass motion. Further, L , \mathcal{J} , and S are the orbital momentum of the center-of-mass motion, the total angular momentum and spin, respectively. We skip for the moment isospin.

The Lippman-Schwinger equation for two-nucleon scattering

For scattering states, the energy is positive, $E > 0$. The Lippman-Schwinger equation (a rewrite of the Schroedinger equation) is an integral equation where we have to deal with the amplitude $R(k, k')$ (reaction matrix, which is the real part of the full complex T -matrix) defined through the integral equation for one partial wave (no coupled-channels)

$$R_l(k, k') = V_l(k, k') + \frac{2}{\pi} \mathcal{P} \int_0^\infty dq q^2 V_l(k, q) \frac{1}{E - q^2/m} R_l(q, k'). \quad (14)$$

For negative energies (bound states) and intermediate states scattering states blocked by occupied states below the Fermi level.

The Lippman-Schwinger equation for two-nucleon scattering

The symbol \mathcal{P} in the previous slide indicates that Cauchy's principal-value prescription is used in order to avoid the singularity arising from the zero of the denominator.

The total kinetic energy of the two incoming particles in the center-of-mass system is

$$E = \frac{k_0^2}{m_n}$$

The Lippman-Schwinger equation for two-nucleon scattering

The matrix $R_l(k, k')$ relates to the the phase shifts through its diagonal elements as

$$R_l(k_0, k_0) = -\frac{\tan \delta_l}{mk_0}. \quad (15)$$

The Lippman-Schwinger equation for two-nucleon scattering

From now on we will drop the subscript l in all equations. In order to solve the Lippman-Schwinger equation in momentum space, we need first to write a function which sets up the mesh points. We need to do that since we are going to approximate an integral through

$$\int_a^b f(x) dx \approx \sum_{i=1}^N w_i f(x_i),$$

where we have fixed N lattice points through the corresponding weights w_i and points x_i . Typically obtained via methods like Gaussian quadrature.

The Lippman-Schwinger equation for two-nucleon scattering

If you use Gauss-Legendre the points are determined for the interval $x_i \in [-1, 1]$. You map these points over to the limits in your integral. You can then use the following mapping

$$k_i = \text{const} \times \tan \left\{ \frac{\pi}{4} (1 + x_i) \right\},$$

and

$$\omega_i = \text{const} \frac{\pi}{4} \frac{w_i}{\cos^2 \left(\frac{\pi}{4} (1 + x_i) \right)}.$$

If you choose units fm^{-1} for k , set $\text{const} = 1$. If you choose to work with MeV, set $\text{const} \sim 200$ ($\hbar c = 197 \text{ MeVfm}$).

The Lippman-Schwinger equation for two-nucleon scattering

The principal value integral is rather tricky to evaluate numerically, mainly since computers have limited precision. We will here use a subtraction trick often used when dealing with singular integrals in numerical calculations. We introduce first the calculus relation

$$\int_{-\infty}^{\infty} \frac{dk}{k - k_0} = 0.$$

It means that the curve $1/(k - k_0)$ has equal and opposite areas on both sides of the singular point k_0 . If we break the integral into one over positive k and one over negative k , a change of variable $k \rightarrow -k$ allows us to rewrite the last equation as

$$\int_0^{\infty} \frac{dk}{k^2 - k_0^2} = 0.$$

The Lippman-Schwinger equation for two-nucleon scattering

We can then express a principal values integral as

$$\mathcal{P} \int_0^{\infty} \frac{f(k) dk}{k^2 - k_0^2} = \int_0^{\infty} \frac{(f(k) - f(k_0)) dk}{k^2 - k_0^2}, \quad (16)$$

where the right-hand side is no longer singular at $k = k_0$, it is proportional to the derivative df/dk , and can be evaluated numerically as any other integral.

The Lippman-Schwinger equation for two-nucleon scattering

We can then use this trick to obtain

$$R(k, k') = V(k, k') + \frac{2}{\pi} \int_0^{\infty} dq \frac{q^2 V(k, q) R(q, k') - k_0^2 V(k, k_0) R(k_0, k')}{(k_0^2 - q^2)/m}. \quad (17)$$

This is the equation to solve numerically in order to calculate the phase shifts. We are interested in obtaining $R(k_0, k_0)$.

The Lippman-Schwinger equation for two-nucleon scattering

How do we proceed?

Using the mesh points k_j and the weights ω_j , we reach

$$R(k, k') = V(k, k') + \frac{2}{\pi} \sum_{j=1}^N \frac{\omega_j k_j^2 V(k, k_j) R(k_j, k')}{(k_0^2 - k_j^2)/m} - \frac{2}{\pi} k_0^2 V(k, k_0) R(k_0, k')$$

The Lippman-Schwinger equation for two-nucleon scattering

This equation contains now the unknowns $R(k_j, k_j)$ (with dimension $N \times N$) and $R(k_0, k_0)$.

We can turn it into an equation with dimension $(N + 1) \times (N + 1)$ with a mesh which contains the original mesh points k_j for $j = 1, N$ and the point which corresponds to the energy k_0 . Consider the latter as the 'observable' point. The mesh points become then k_j for $j = 1, n$ and $k_{N+1} = k_0$.

With these new mesh points we define the matrix

$$A_{i,j} = \delta_{i,j} - V(k_i, k_j) u_j, \quad (18)$$

The Lippman-Schwinger equation for two-nucleon scattering

where δ is the Kronecker δ and

$$u_j = \frac{2}{\pi} \frac{\omega_j k_j^2}{(k_0^2 - k_j^2)/m} \quad j = 1, N$$

and

$$u_{N+1} = -\frac{2}{\pi} \sum_{j=1}^N \frac{k_0^2 \omega_j}{(k_0^2 - k_j^2)/m}.$$

The Lippman-Schwinger equation for two-nucleon scattering

The first task is then to set up the matrix A for a given k_0 . This is an $(N+1) \times (N+1)$ matrix. It can be convenient to have an outer loop which runs over the chosen observable values for the energy k_0^2/m . Note that all mesh points k_j for $j = 1, N$ must be different from k_0 . Note also that $V(k_i, k_j)$ is an $(N+1) \times (N+1)$ matrix. With the matrix A we can rewrite the problem as a matrix problem of dimension $(N+1) \times (N+1)$. All matrices R , A and V have this dimension and we get

$$A_{i,j} R_{i,j} = V_{i,j},$$

or just

$$AR = V.$$

The Lippman-Schwinger equation for two-nucleon scattering

Since you already have defined A and V (these are stored as $(N+1) \times (N+1)$ matrices) The final equation involves only the unknown R . We obtain it by matrix inversion, i.e.,

$$R = A^{-1}V. \quad (19)$$

Thus, to obtain R , you will need to set up the matrices A and V and invert the matrix A . With the inverse A^{-1} , perform a matrix multiplication with V results in R .

The Lippman-Schwinger equation for two-nucleon scattering

With R you can then evaluate the phase shifts by noting that

$$R(k_{N+1}, k_{N+1}) = R(k_0, k_0) = -\frac{\tan \delta}{mk_0},$$

where δ are the phase shifts.

The Lippman-Schwinger equation for two-nucleon scattering

For elastic scattering, the scattering potential can only change the outgoing spherical wave function up to a phase. In the asymptotic limit, far away from the scattering potential, we get for the spherical bessel function

$$j_l(kr) \xrightarrow{r \gg 1} \frac{\sin(kr - l\pi/2)}{kr} = \frac{1}{2ik} \left(\frac{e^{i(kr - l\pi/2)}}{r} - \frac{e^{-i(kr - l\pi/2)}}{r} \right)$$

The outgoing wave will change by a phase shift δ_l , from which we can define the S-matrix $S_l(k) = e^{2i\delta_l(k)}$. Thus, we have

$$\frac{e^{i(kr - l\pi/2)}}{r} \xrightarrow{\text{phasechange}} \frac{S_l(k) e^{i(kr - l\pi/2)}}{r}$$

The Lippman-Schwinger equation for two-nucleon scattering

The solution to the Schrodinger equation for a spherically symmetric potential, will have the form

$$\psi_k(r) = e^{ikr} + f(\theta) \frac{e^{ikr}}{r}$$

where $f(\theta)$ is the scattering amplitude, and related to the differential cross section as

$$\frac{d\sigma}{d\Omega} = |f(\theta)|^2$$

Using the expansion of a plane wave in spherical waves, we can relate the scattering amplitude $f(\theta)$ with the partial wave phase shifts δ_l by identifying the outgoing wave

$$\psi_k(r) = e^{ikr} + \left[\frac{1}{2ik} \sum_l i^l (2l+1) (S_l(k) - 1) P_l(\cos(\theta)) e^{-il\pi/2} \right] \frac{e^{ikr}}{r}$$

which can be simplified further by cancelling i^l with $e^{-il\pi/2}$

The Lippman-Schwinger equation for two-nucleon scattering

We have

$$\psi_k(r) = e^{ikr} + f(\theta) \frac{e^{ikr}}{r}$$

with

$$f(\theta) = \sum_l (2l+1) f_l(\theta) P_l(\cos(\theta))$$

where the partial wave scattering amplitude is given by

$$f_l(\theta) = \frac{1}{k} \frac{(S_l(k) - 1)}{2i} = \frac{1}{k} \sin \delta_l(k) e^{i\delta_l(k)}$$

With Eulers formula for the cotangent, this can also be written as

$$f_l(\theta) = \frac{1}{k} \frac{1}{\cot \delta_l(k) - i}$$

The Lippman-Schwinger equation for two-nucleon scattering

Interpretation of phase shifts

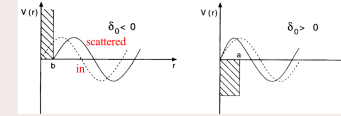


Figure: Examples of negative and positive phase shifts for repulsive and attractive potentials, respectively.

The Lippman-Schwinger equation for two-nucleon scattering

The integrated cross section is given by

$$\begin{aligned} \sigma &= 2\pi \int_0^\pi |f(\theta)|^2 \sin \theta d\theta \\ &= 2\pi \sum_l \left| \frac{(2l+1)}{k} \sin(\delta_l) \right|^2 \int_0^\pi (P_l(\cos(\theta)))^2 \sin(\theta) d\theta \\ &= \frac{4\pi}{k^2} \sum_l (2l+1) \sin^2 \delta_l(k) = 4\pi \sum_l (2l+1) |f_l(\theta)|^2, \end{aligned}$$

where the orthogonality of the Legendre polynomials was used to evaluate the last integral

$$\int_0^\pi P_l(\cos \theta)^2 \sin \theta d\theta = \frac{2}{2l+1}$$

Thus, the **total** cross section is the sum of the partial-wave cross sections. Note that the differential cross section contains cross-terms from different partial waves. The integral over the full

The Lippman-Schwinger equation for two-nucleon scattering

At low energy, $k \rightarrow 0$, S-waves are most important. In this region we can define the scattering length a and the effective range r . The S-wave scattering amplitude is given by

$$f_l(\theta) = \frac{1}{k} \frac{1}{\cot \delta_l(k) - i}$$

Taking the limit $k \rightarrow 0$, gives us the expansion

$$k \cot \delta_0 = -\frac{1}{a} + \frac{1}{2} r_0 k^2 + \dots$$

Thus the low energy cross section is given by

$$\sigma = 4\pi a^2$$

If the system contains a bound state, the scattering length will become positive (neutron-proton in 3S_1). For the 1S_0 wave, the scattering length is negative and large. This indicates that the wave function of the system is at the verge of turning over to get a node,

The Lippman-Schwinger equation for two-nucleon scattering

Low energy scattering length

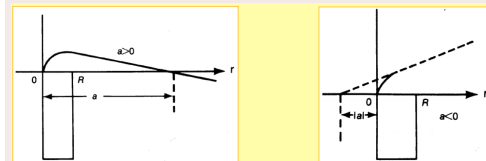
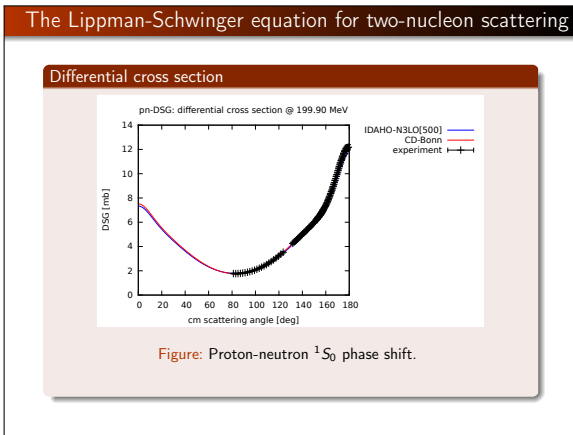
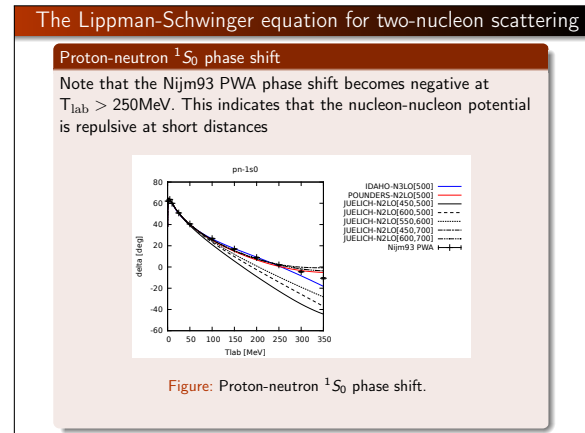
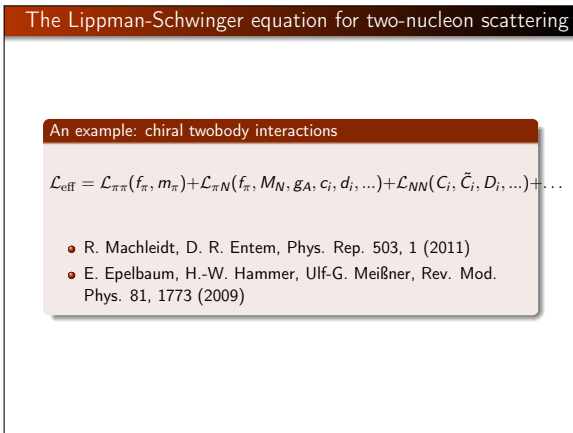
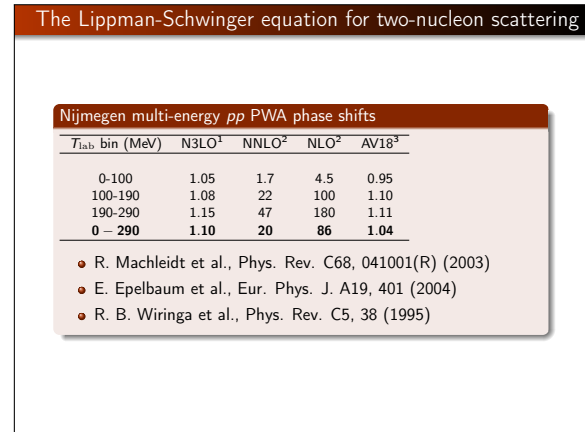
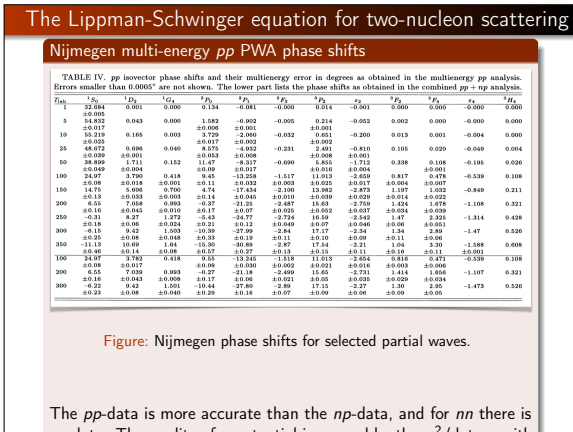


Figure: Examples of scattering lengths.

The Lippman-Schwinger equation for two-nucleon scattering

It is important to realize that the phase shifts themselves are not observables. The measurable scattering quantity is the cross section, or the differential cross section. The partial wave phase shifts can be thought of as a parameterization of the (experimental) cross sections. The phase shifts provide insights into the physics of partial wave projected nuclear interactions, and are thus important quantities to know.

The nucleon-nucleon differential cross section have been measured at almost all energies up to the pion production threshold (290 MeV in the Lab frame), and this experimental data base is what provides us with the constraints on our nuclear interaction models. In order to pin down the unknown coupling constants of the theory, a statistical optimization with respect to cross sections need to be carried out. This is how we constrain the nucleon-nucleon interaction in practice!



a) List all allowed according to the Pauli principle partial waves with isospin T , their projection T_z , spin S , orbital angular momentum l and total spin J for $J \leq 3$. Use the standard spectroscopic notation $^{2S+1}L_J$ to label different partial waves. A proton-proton state has $T_z = -1$, a proton-neutron state has $T_z = 0$ and a neutron-neutron state has $T_z = 1$.

- a) Find the closed form expression for the spin-orbit force. Show that the spin-orbit force $\mathbf{L}\mathbf{S}$ gives a zero contribution for S -waves (orbital angular momentum $l = 0$). What is the value of the spin-orbit force for spin-singlet states ($S = 0$)?
- b) Find thereafter the expectation value of $\sigma_1 \cdot \sigma_2$, where σ_i are so-called Pauli matrices.
- c) Add thereafter isospin and find the expectation value of $\sigma_1 \cdot \sigma_2 \tau_1 \cdot \tau_2$, where τ_i are also so-called Pauli matrices. List all the cases with $S = 0, 1$ and $T = 0, 1$.

A simple parametrization of the nucleon-nucleon force is given by what is called the V_8 potential model, where we have kept eight different operators. These operators contain a central force, a spin-orbit force, a spin-spin force and a tensor force. Several features of the nuclei can be explained in terms of these four components. Without the Pauli matrices for isospin the final form of such an interaction model results in the following form:

$$V(r) = \left\{ C_c + C_\sigma \sigma_1 \cdot \sigma_2 + C_T \left(1 + \frac{3}{m_\alpha r} + \frac{3}{(m_\alpha r)^2} \right) S_{12}(\hat{r}) + C_{SL} \left(\frac{1}{m_\alpha r} + \frac{1}{(m_\alpha r)^2} \right) \mathbf{L} \cdot \mathbf{S} \right\} \frac{e^{-m_\alpha r}}{m_\alpha r}$$

where m_α is the mass of the relevant meson and S_{12} is the familiar tensor term. The various coefficients C_i are normally fitted so that the potential reproduces experimental scattering cross sections. By adding terms which include the isospin Pauli matrices results in an interaction model with eight operators.

The expectation value of the tensor operator is non-zero only for $S = 1$. We will show this in a forthcoming lecture, after that we

The aim here is to develop a program which solves the Lippman-Schwinger equation for a simple parametrization for the 1S_0 partial wave. This partial wave is given by a central force only and is parametrized in coordinate space as

$$V(r) = V_a \frac{e^{-ax}}{x} + V_b \frac{e^{-bx}}{x} + V_c \frac{e^{-cx}}{x}$$

with $x = \mu r$, $\mu = 0.7$ fm (the inverse of the pion mass), $V_a = -10.463$ MeV and $a = 1$, $V_b = -1650.6$ MeV and $b = 4$ and $V_c = 6484.3$ MeV and $c = 7$.

- a) Find an analytical expression for the Fourier-Bessel transform (Hankel transform) to momentum space for $l = 0$ using

$$\langle k | V_l | k' \rangle = \int j_l(kr) V(r) j_l(k'r) r^2 dr.$$

- b) Write a small program which calculates the latter expression and use this potential to compute the T -matrix at positive energies for $l = 0$. Compare your results to those obtained with a box potential given by

$$V(r) = \begin{cases} -V_0 & r < R_0 \\ 0 & r > R_0 \end{cases}$$